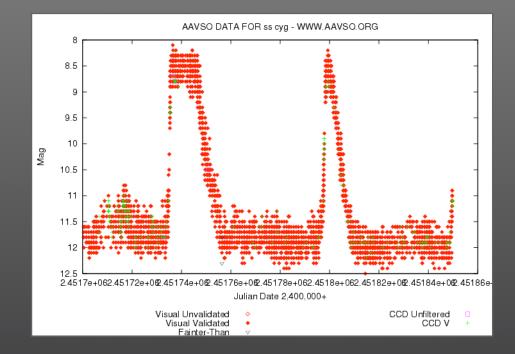
A New Look At SS Cyg

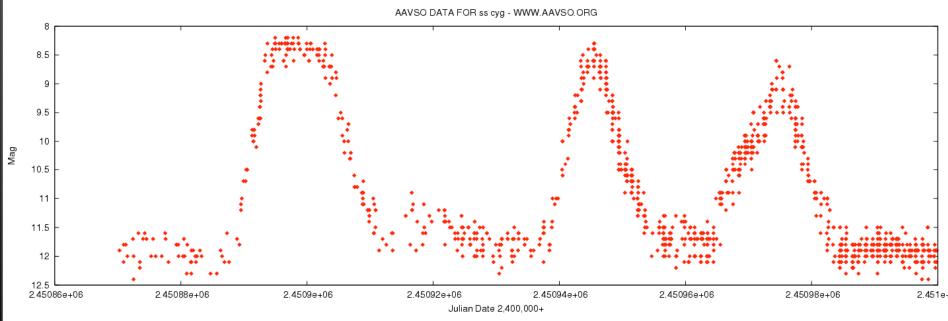


Aaron Price, AAVSO 95th Spring Meeting of the AAVSO Rockford, IL

SS Cyg Outbursts

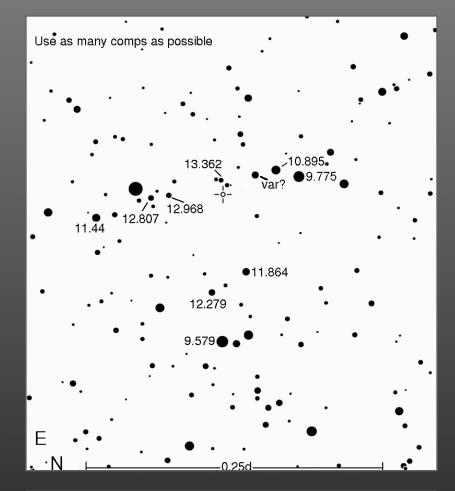
- Long/Wide (>I2d)
- Short/Narrow (<I2d)
- Anomalous
- Cannizo & Mattei (1998)
- Fast (82%: ~2.3 mags/day)
- Slow (~0.5 mags/day)
- The "Glitch"
- Sampling





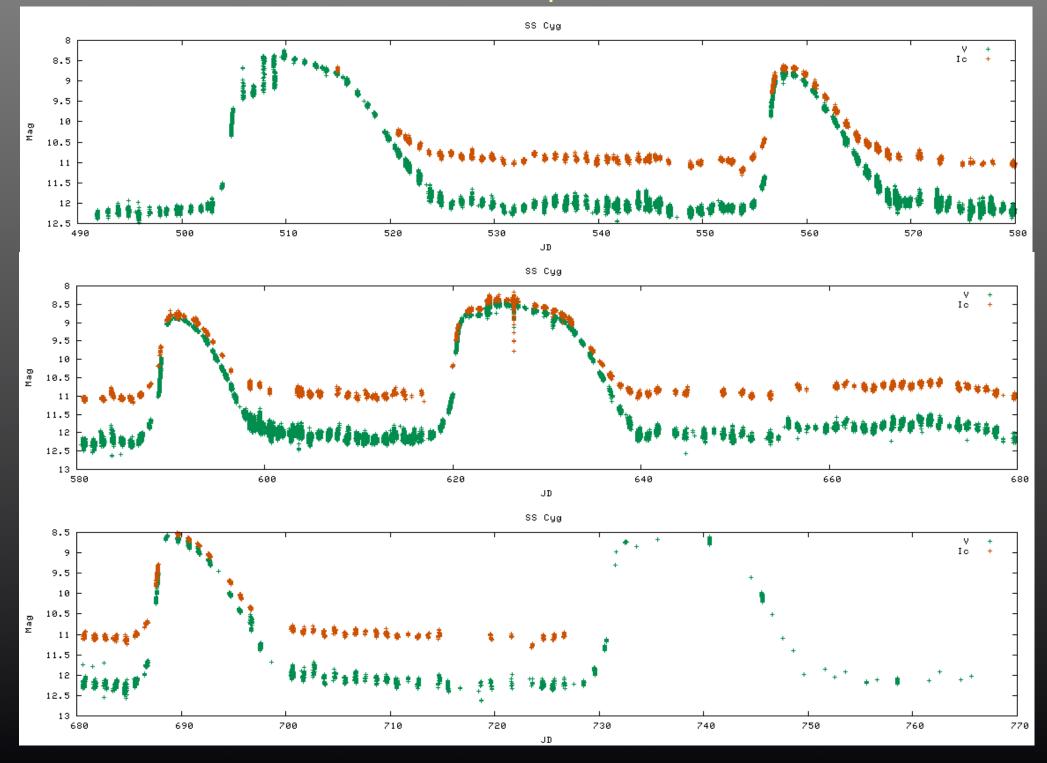
At a Glance

- Primary Goal: Test for outburst precursors
- Secondary Goals: Search for the "glitch" and any other recurring behavior; Higher resolution tests of previously published results
- Tertiary Goal: Create a high-quality, intensive data set for use by professional researchers
- Method: High precision CCD data for as long as possible, each night.
- Preparation and practice in April, 2005
- Full campaign May I Oct. I, 2005
- Volunteer follow up Oct. I Jan. 29, 2006
- Ic began May 29, 2005



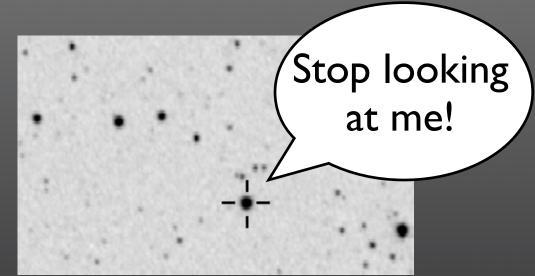
v	B-V	V-R	R-lc	
9.775	0.379	NA	NA	
10.895	0.522	0.312	0.270	
11.44	0.600	NA	NA	
11.864	0.620	0.358	0.328	

Let there be photons...



And it was good...

- 95, 720 V Observations
- 11,467 lc Observations
- 274 days of coverage
- ~134,496 minutes of observing time
- 2,241.6 hours
- 93.4 days
- 3,672 cups o' joe
- 165 spousal permission units
- 13 bouquets of flowers
- I broken server/light curve generator
- V = 85,392m/1,423.2h/59.3d
- lc = 49,104m/818.4h/34.1d

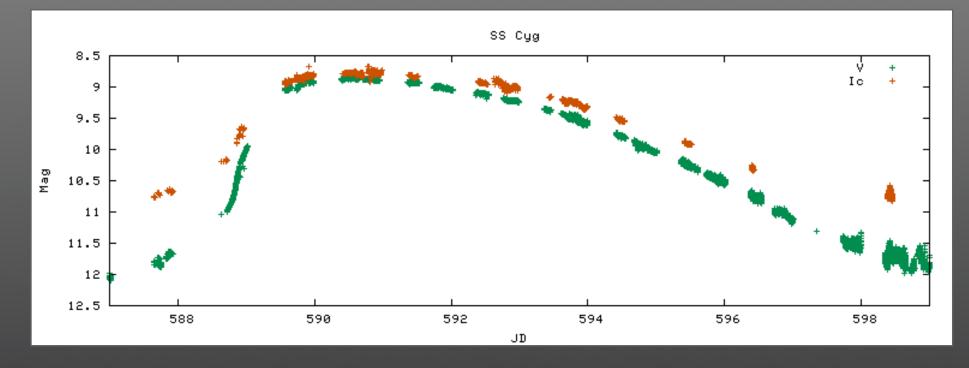


The 2005 SS Cyg Hall of Fame

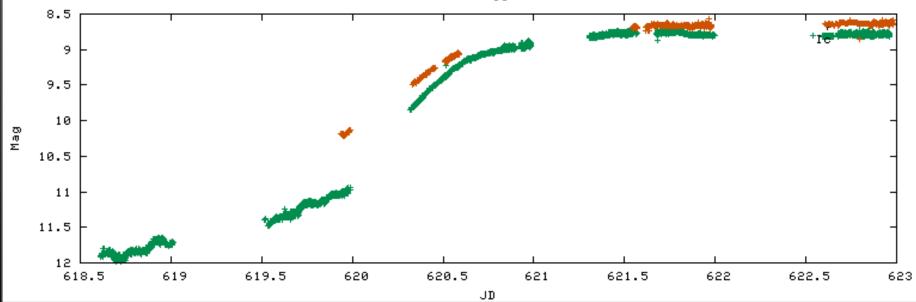
Canada, Belgium, England, United States

JM	Robert James			
PVA	Vance Petriew			
HUZ	Rick Huziak			
SDB	Donn Starkey			
MXL	Richard Miles			
GFB	Bill Goff			
VMT	Tonny Vanmunster			
TDY	Dave Tandy			
BKL	John Blackwell			
BDG	David Boyd			
WGR	Gary Walker			
BXS	Steve Brady			
KMP	Mike Koppelman			
CTX	Tim Crawford			
JJI	Jim Jones			
PJT	Jennifer Petriew			
SGRA	Graham Salmon			
GBL	Bruce Gary			

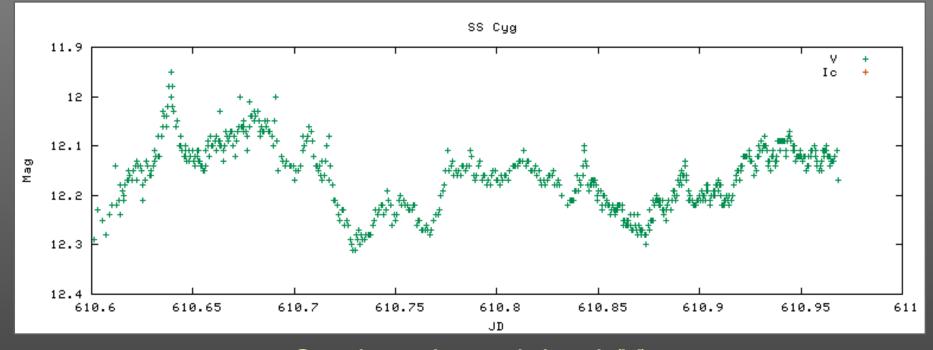
Close Ups



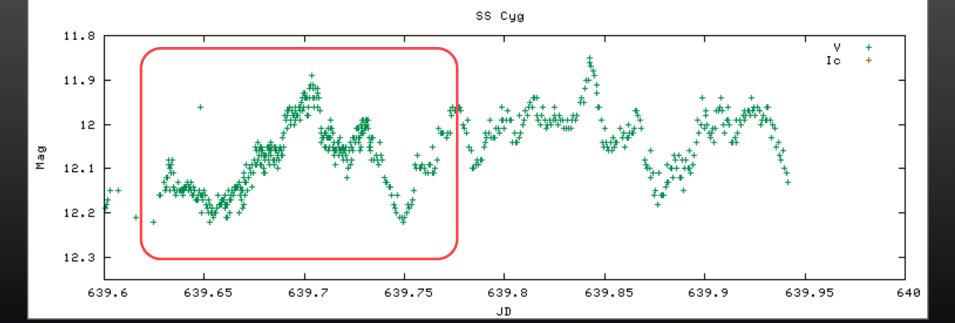
SS Cyg



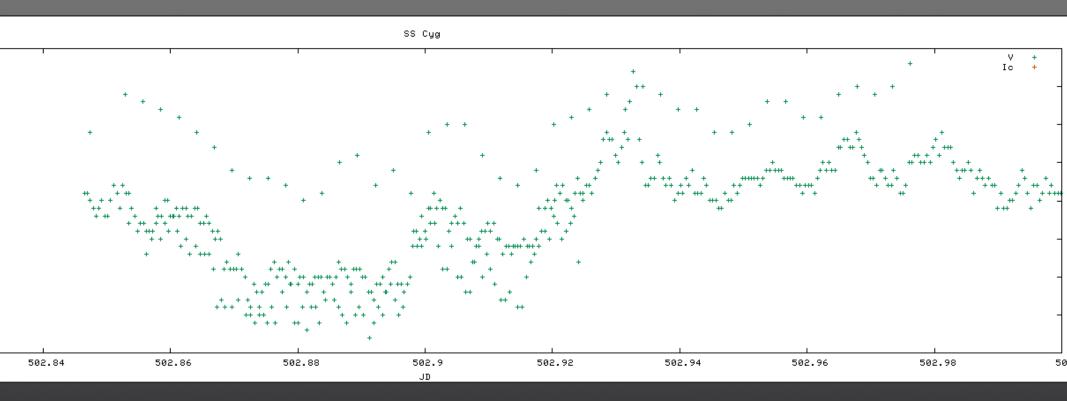
4 Observers (V)



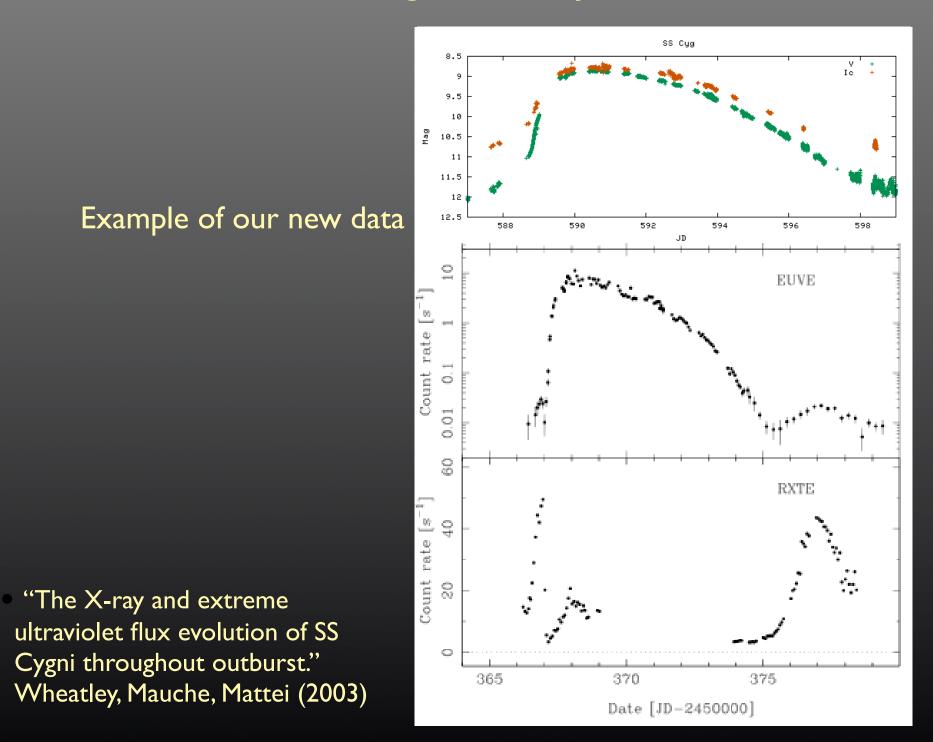
Good overlap and detail (V)



Comp Star & Chart Problems



Multiwavelength Comparison



Robert L. Hill (Ball State), AAS 205th meeting (2005)

- Subset of AAVSO visual data
- Flux converted, Binned, averaged, normalized to preoutburst magnitude
 SS Cyg brightens by 2% I day before outburst

Outburst Energies and Precursors to Outbursts for SS Cyg from AAVSO Data

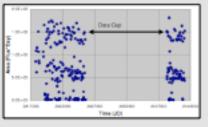
Method: Outburst Energies

I obtained two data sets for SS Cyg. I was given the first data set by my adviser. Dr. Ron Kaltchuck. I downloaded the second data set from the AAVSO website.

The magnitudes were converted to fluxes using F(L) = 10^{-(LA)/dead magnitude(-8.3)}.

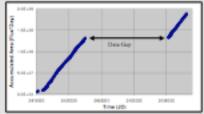
The area of each outburst on a flux vs. time graph was determined using the Trapezoidal rule. The outburst area should be proportional to the energy released during the outburst.

Outburst Energies



Implication Three types of outbursts occur. They do not occur in a very regular pattern.

Accumulated Outburst Energies



Implication The accumulation of outburst energy is linear over time. This implies that the average mass transfer rate for a year is the same as the mass transfer rate for another year.

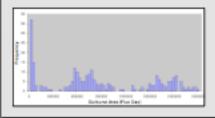
Three types of outbursts occur

Two outburst types might be

Implication

bimodal.

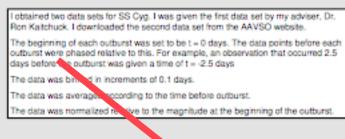
Outburst Energy Histogram



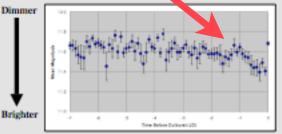
Author: Robert Hill (roberthill58@yahoo.com, rihill@bsu.edu) Department of Physics and Astronomy Ball State University

Co-Author: E.O. Waggen AAVSO

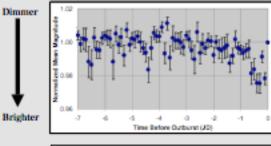
Method: Outburst Precursors



Mean Magnitude Before Outburst



Normalized Mean Magnitude Before Outburst

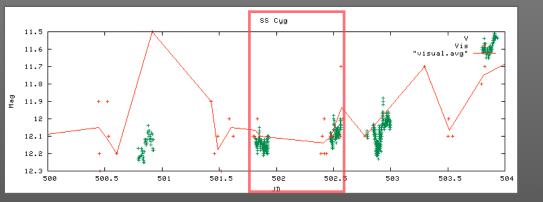


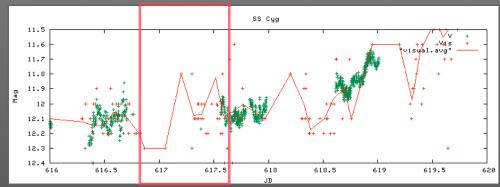
Implication

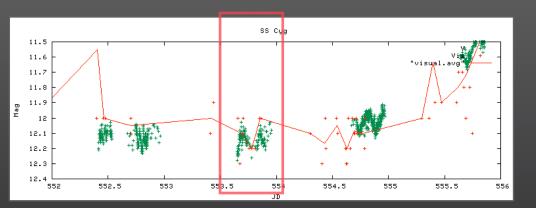
SS Cyg appears to brighten around 2% on the average on the day before going into outburst.

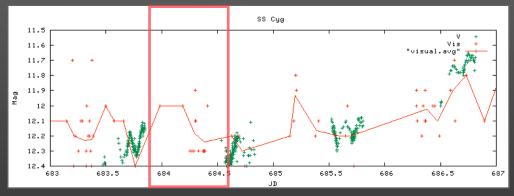


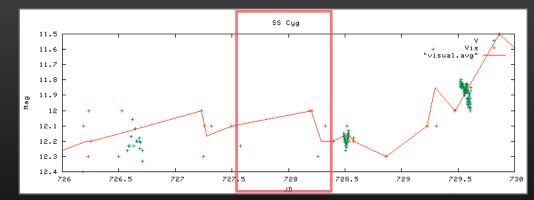
4 Days Around Outburst Onset



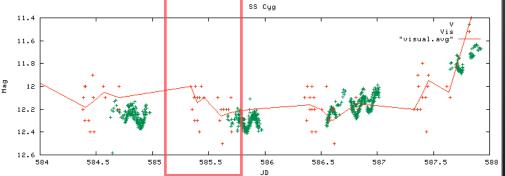








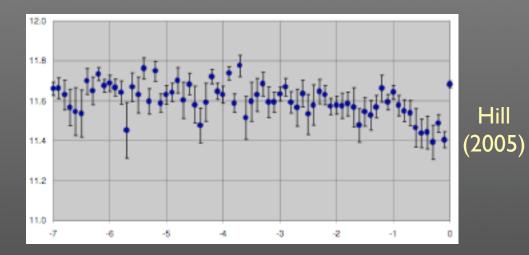
Visual+0.1d avg

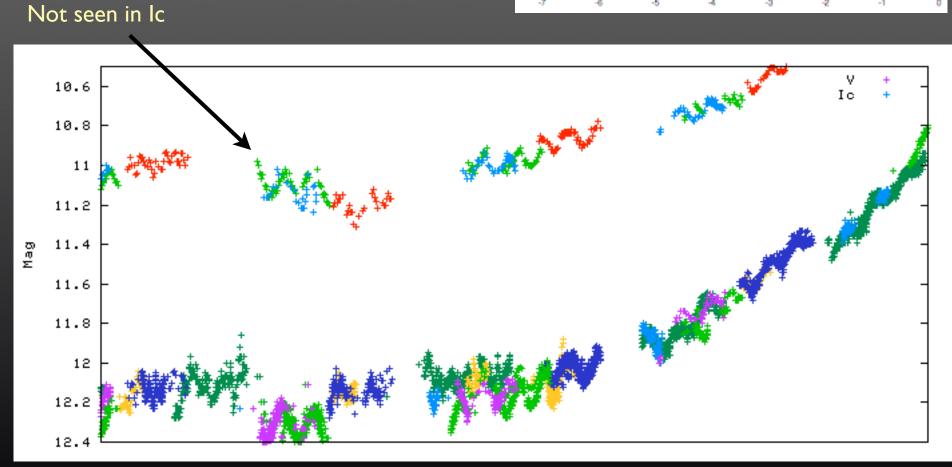


JV

Combined Data

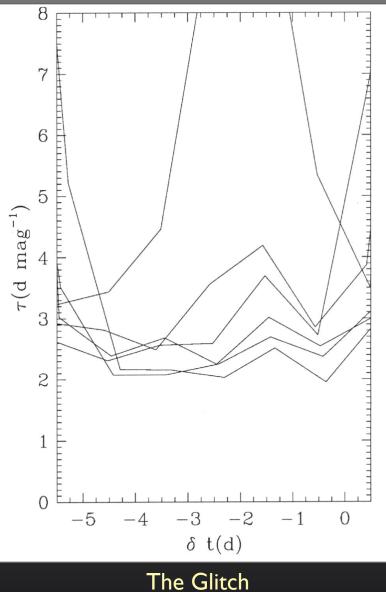
- Flickering amplitude is too great to determine: need long term monitoring to average out
- Interesting feature in 2 data sets ~I-2 days before outburst onset
- Phasing of the light curve?





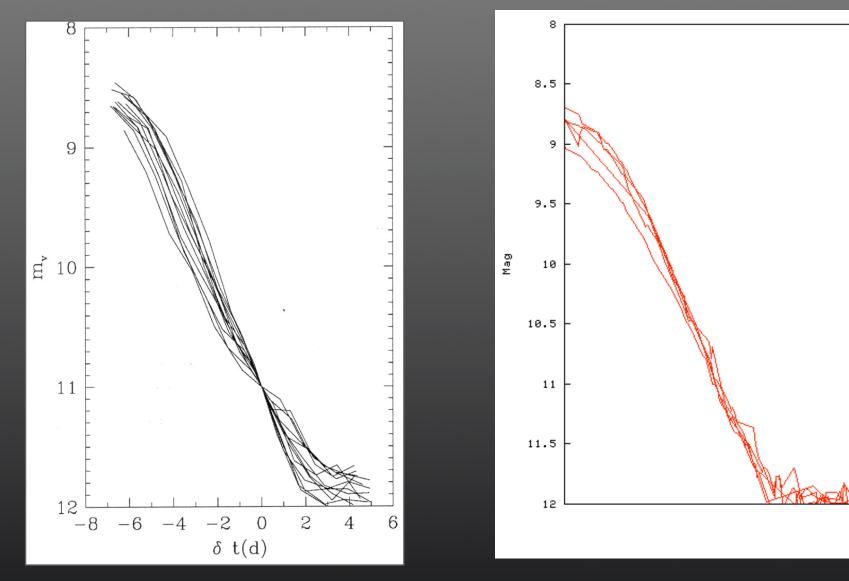
The Cannizo Glitch

- Decays are largely exponential
- Except for a "glitch" around 2/3 through (2 days before the end)
- The larger the time of decay, the larger the glitch
- 20% 300% increase in t_{decay} for ~ I day



(Fig. 7 from Cannizo & Mattei (1998))

Looking For the Glitch

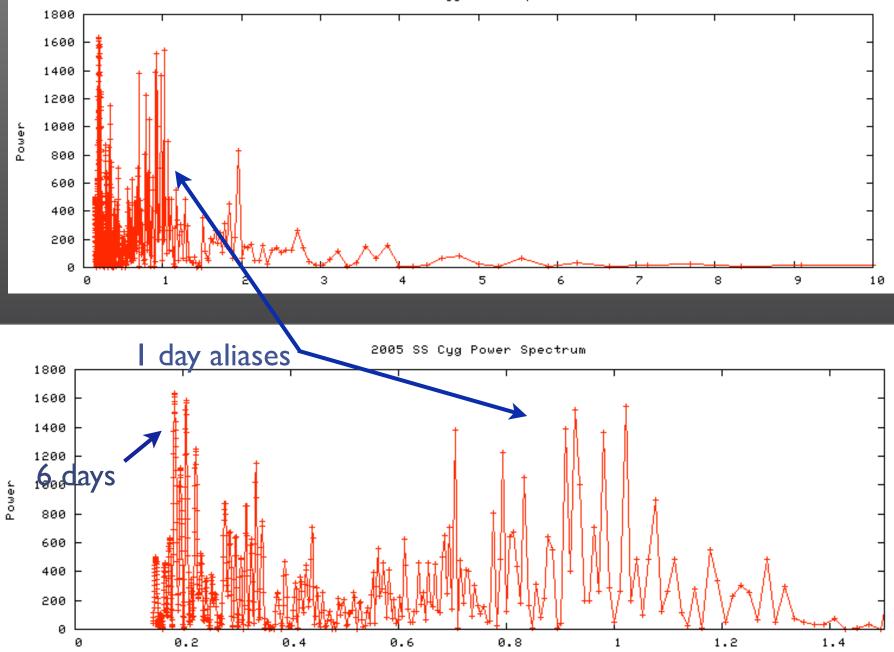


The Glitch, correlated at Mv=11 (Fig. 8 from Cannizo & Mattei (1998)) Our data + 1d means, correlated at Mv=11

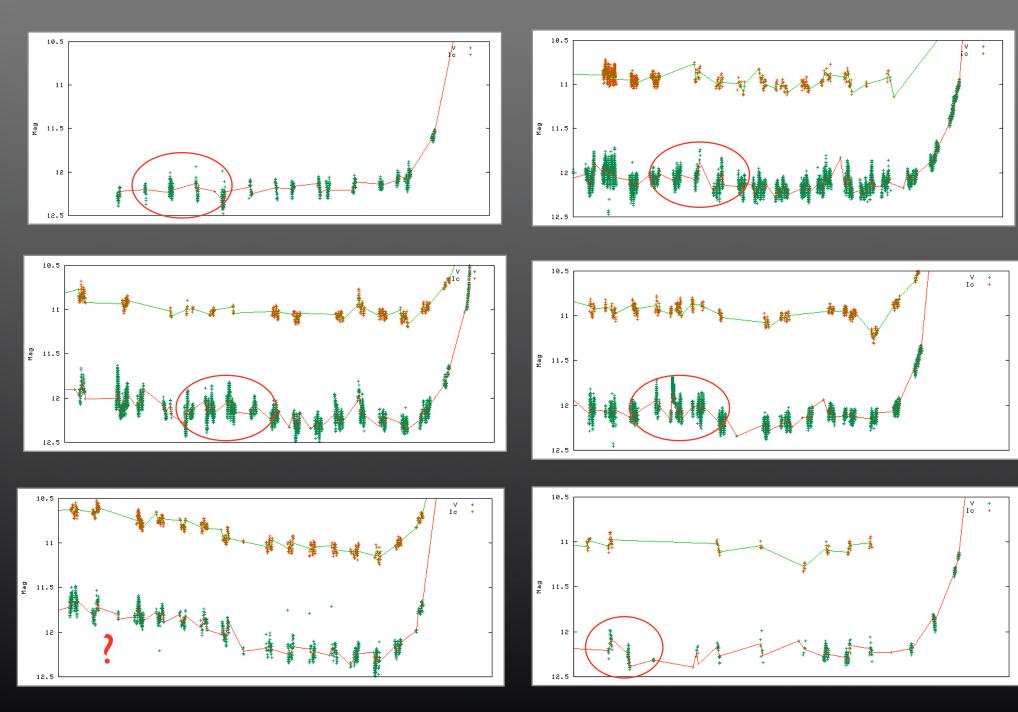


Periodicity

2005 SS Cyg Power Spectrum

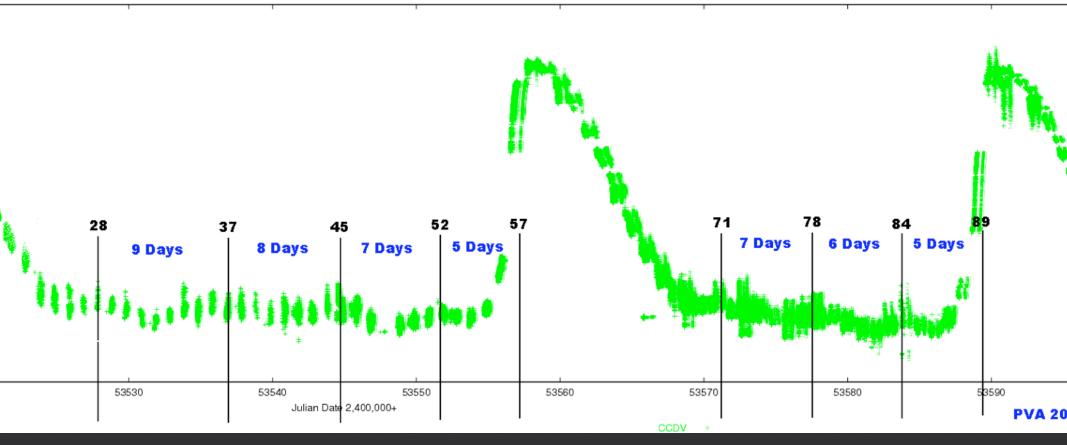


Precursor Humps



Quiescent Modulation

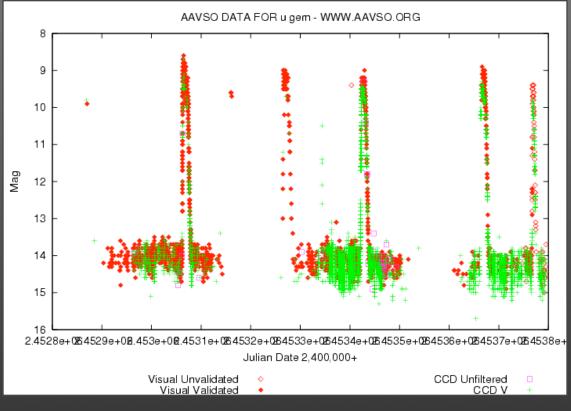
AAVSO UNVALIDATED DATA FOR SS CYG - WWW.AAVSO.ORG



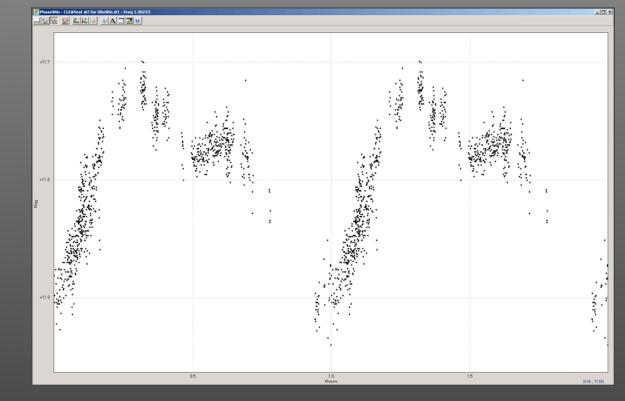
Graph by Vance Petriew (PVA)

Side Benefits

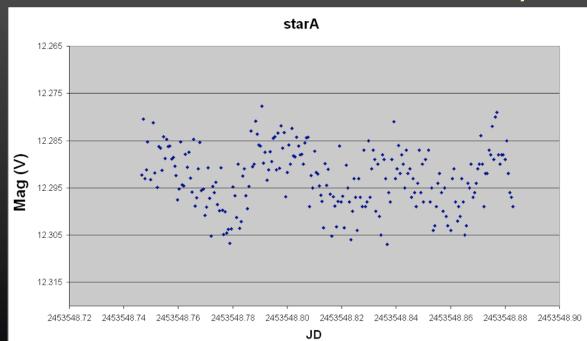
- HUZ and PVA are conducting searches for new field variables
- I confirmed suspected variable star (a comp star!)
- I newly discovered variable star (RRab)
- 3 suspected new variables
- Observer experience with time series
- Staff experience with large-scale data processing



Robert James (JM) U Gem CCDV data



New variables courtesy Rick Huziak (HUZ)



Now What?

- Rest!
- Archival visual data statistical analysis
- Paper in summer?
- SMOOTCH!
- U Gem in B and V?
- More SS Cyg: nightly snapshots indefinitely



Some could care less about SS Cyg



CCD & Visual Animation by Vance Petriew From the Light Curve Generator

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